THE CIRCUMSTELLAR ENVIRONMENT OF EVOLVED STARS
AS REVEALED BY STUDIES OF CIRCUMSTELLAR WATER
MASERS

BY
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A Dissertation submitted to the Graduate School
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"The Circumstellar Environment of Evolved Stars as Revealed by Proper Motion Studies of Circumstellar Water Masers," a dissertation prepared by Kevin Boyd Marvel in partial fulfillment of the requirements for the degree, Doctor of Philosophy, has been approved and accepted by the following:

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“That this book has its faults, no one can doubt,

Although the Author could not find them out.

The faults you find, good Reader, please to mend,

Your comments to the Author kindly send.”

*Kitchiner’s The Economy of the Eyes.*—Part II.

When I first decided to become an astronomer, I gave up a promising career in marine biology. Of course, being only twelve years old at the time it wasn’t such a great loss. I decided to pursue astronomy simply because my parents, June and Barry Marvel, purchased a cheap 1.5” refracting telescope and gave it to me for Christmas. I took it outside and pointed it towards the brightest object in the sky (Venus at the time) and found, to my amazement, that this little tube took a small, point-like image on the sky and displayed a small crescent shaped object in the eyepiece. I was just amazed. I had no idea what I was looking at, so I ran inside and managed to find an old *Zim’s Field Guide to the Stars*, which explained that Venus went through phases and that this was a major piece of evidence for the Copernican model of the Solar System. Well, I was hooked. Since that time, no matter what crazy ideas I came up with (like getting up at 3:00 a.m. to see meteors, traveling through scary third-world countries to see an eclipse, or going to school for eleven years), my parents have
supported me completely. They provided me with an early exposure to all forms of human endeavor from art to construction, an absolute must for a scientist. To them I owe my college education at a world-renowned astronomical center, which formed the basis of my skills allowing the completion of this dissertation. To them I owe everything, and little else can be said except that I thank you both from the bottom of my heart for helping me fulfill one of my life’s goals. I hope this makes up for setting the backyard on fire that one time.

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When I finally got to college at the University of Arizona, I met a group of individuals with whom I lived, worked, played and learned. Luke Keller, Babar Ali and Tad Adair: You know who you are and what you did… thanks bunches. Tad helped out with some statistical analysis and figure creation for this dissertation. All the other science nerds, who prided themselves on not being astronomers, thanks to you as well for keeping my mind open to other approaches: Jen Ahearn, Keith Mulvihill, Jenny Vuturo, Heather Merbs, Teri Suzuki, Andy Young, Brett McDaniel, and anyone else I forgot, thank you for always being around… we sure had fun didn’t we?

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At the University of Arizona, nearly all my astronomy classes were taught by one individual, Dr. Thomas Swihart. As an educator and advisor he was without equal. He knew the value of estimation and detail. I regret that he passed away prior to the completion of my doctorate, but I think he never had any doubts I could finish, even when I did. He gave sage advice to all his students, and I will never forget the first day of Astronomy 271 when he lectured for a whole hour on how difficult the class would be and that all the students should drop now. I think only a few did, but his point was that you should only attempt things if you want to work hard and struggle. Thank you, Dr. Swihart, and may you rest in peace.

I managed to get into graduate school because one man called me on the phone and wanted to know what was up with my GPA. After explaining my situation (I guess to his satisfaction), Dr. Jack Burns let me into his department as a trial experiment. Along with my eight classmates, I began in 1990 to learn more at a faster rate than I ever had before. I stumbled a few times along the way, but Jack was always willing to give me another chance to prove myself. I appreciate his guidance and knowledge...not to mention his often swift kicks in my rear-end. I
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Sigma Zi

American Association of Variable Star Observers

MENSA

FIELDS OF STUDY

Major Field: Astronomy

Very Long Baseline Interferometry

Astrophysical Masers

AGB Stars
ABSTRACT

THE CIRCMSTEMellar ENVIRONMENT OF EVOLVED STARS AS
REVEALED BY STUDIES OF CIRCMSTEMellar WATER MASERS

BY

KEVIN BOYD MARVEL, B.S., B.S., M.S.

Doctor of Philosophy in Astronomy

New Mexico State University

Las Cruces, New Mexico, 1996

Dr. Jack O. Burns, Chair

The dissertation presents the results of a multi-epoch very long baseline interferometric study of water masers located in the extended atmospheres of evolved stars. The research was performed using the Very Long Baseline Array and Very Large Array of the National Radio Astronomy Observatory. Optical monitoring of the stars was provided by the American Association of Variable Star Observers, the Variable Star Network and Dr. Bill Neely of the NF/ Observatory.

Water masers are found to exist in a region where a population inversion of
the rotational transition at 22 GHz can be maintained by collisional pumping.

The masers are identified as individual pockets of gas, which have good velocity coherence and may be imaged using radio interferometry. Stellar winds are initiated in these sources by dust formation and acceleration of the gas through momentum coupling. The typical wind speeds in the region of the water masers are 10 to 20 km/s. The water masers are followed through several epochs of observation and exhibit proper motions consistent with the assumed source distance and the measured outflow velocity in the water maser region. Estimates of the distance to the sources using statistical approximation are in agreement with the currently accepted distances to the stars. A detailed kinematic model is used to describe the flow motions of the gas in the maser region. The regions are found to be complex and not well modeled by uniform radial outflow, radial outflow with rotation, or radial outflow with acceleration. The reasons for this are explored and include anisotropic velocity fields induced through non-uniform dust formation near the star and incomplete sampling of the outflow due to a lack of detected masers. Possibilities for future work in the subject are described and include more sophisticated modeling, more sensitive observations, and analysis of other maser species.
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